An Ultracold Neutron Source for TRIUMF

Jeff Martin (U. Winnipeg)

for Lothar Buchmann, Chuck Davis, Michael Gericke, Bob Golub, Akira Konaka, Larry Lee, Yasuhiro Masuda, Ania Micherdzinska, Shelley Page, Des Ramsay, Wim van Oers

(KEK, TRIUMF, NCSU, U. Manitoba, U. Winnipeg)

LABORATOIRE NATIONAL CANADIEN POUR LA RECHERCHE EN PHYSIQUE NUCLÉAIRE ET EN PHYSIQUE DES PARTICULES

Propriété d'un consortium d'universités canadiennes, géré en co-entreprise à partir d'une contribution administrée par le Conseil national de recherches Canada

Outline

- 1. Introduction to Ultracold Neutrons (UCN)
- 2. UCN physics experiments
- 3. Source work at RCNP for TRIUMF
- 4. Plans for TRIUMF experiments

Ultracold Neutrons

- UCN are neutrons that are moving so slowly that they are totally reflected from surfaces of materials.
- So, they can be confined in material bottles for long periods of time.
- Typical parameters:
 - velocity < 8 m/s
 - temperature < 4 mK
 - kinetic energy < 300 neV
- Interactions:
 - gravity: V=mgh (h < 3 m)
 - weak interaction (allows UCN to decay)
 - magnetic fields: V=-μ•B (100% polarization)
 - strong interaction
- Experiments at UCN sources are chronically limited by UCN density. TRIUMF has the potential to be a world leader in this regard.

UCN Source at TRIUMF would be a world-class facility

	Source type	E_c and τ_s	UCN density ρ _{υςΝ} (UCN/cm³)
TRIUMF 5 kW _{av} proton	0.8K He-II	$E_c = 210 \text{ neV} \ \tau_s = 150 \text{ s}$	1.8 x 10 ⁴ at experimental port
Grenoble 60MW reactor	0.5K He-II	$E_c = 250 \text{ neV} \ \tau_s = 150 \text{ s}$	1000 in He-II
SNS cold neutron beam	0.3K He-II	$E_c = 134 \text{ neV} \tau_s = 500 \text{ s}$	430 in He-II
Munich 20MW reactor	SD ₂	$E_c = 250 \text{ neV}$	10⁴ in source
North Carolina 1 MW reactor	SD_2	E _c = 335 neV	1300 in source
PSI 12 kW _{av} proton	SD_2	$E_c = 250 \text{ neV} \ \tau_s = 888 \text{ s}$	2000 in source
Los Alamos 2.4 kW _{av} proton	SD_2	$E_c = 250 \text{ neV} \ \tau_s = 2.6 \text{ s}$	120 in source

The Most Successful UCN Physics Experiments to Date

- n-EDM
- n-Lifetime
- Quantization of neutron energy levels in Earth's gravitation field

all were done at ILL, Grenoble, France

all were limited by the density of UCN that could be achieved

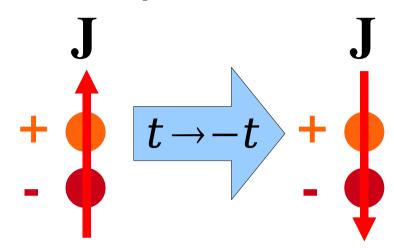
Future UCN Experiments

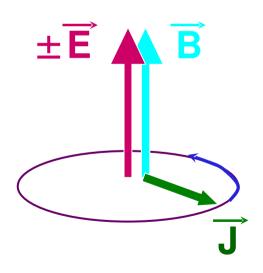
- n-EDM
 - in vacuo (PSI)
 - in 4He/3He (SNS)
 - ILL
 - KEK/JPARC
- n-Lifetime
 - gravito-magnetic bottle (LANL)
 - magnetic bottle in 4He (SNS)
 - multiple ILL proposals

- Gravity Levels
 - ILL
- β -Decay Correlations (A)
 - UCNA (LANL)
- r-process, free neutron target
- nn oscillations
- mirror neutrons
- surface physics
- physics of UCN production
 - e.g. O2

Neutron Electric Dipole Moment (n-EDM)

- Existence of EDM implies violation of Time Reversal Invariance
- CPT Theorem then implies violation of CP conservation

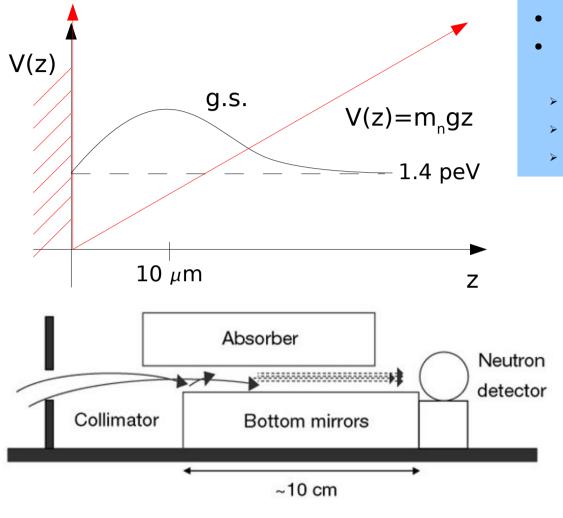




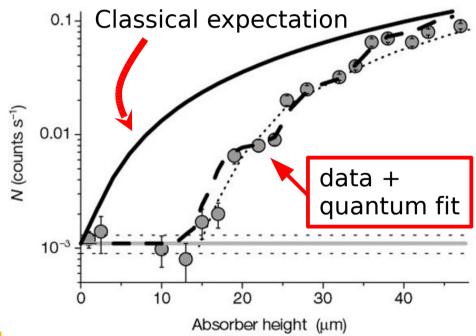
$$h \nu = 2 \mu_n B \pm 2 d_e E$$

- Present Exp. Limit $< 0.63 \times 10^{-25}$ e-cm
- Standard Model value: 10⁻³¹ e-cm
- Supersymmetry or Multi-Higgs models can give 10⁵xSM
- Significant discovery potential with new high sensitivity n-EDM experiment

UCN Quantum States in Gravity



- Confine UCN in 1D by gravity
- Experimental results have been used to place limits on
 - 10 um scale modifications to gravity
 - extra dimensions
 - axions

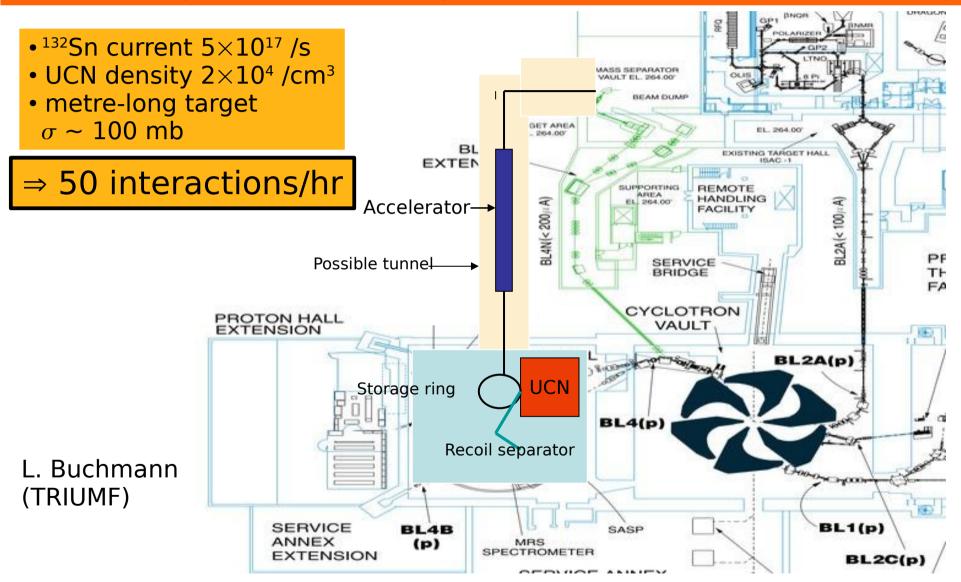


Further experiments:

- Bottle the UCN to increase time the UCN is contact with the mirror.
- Excite resonant transitions between quantum states.
- Increase purity of states by preselection.
- Goal: improve precision on energy of state and hence increase sensitivity to modifications to gravity.

Measuring (n,y) cross sections of the r-process

radioactive species stored in ring interact with free neutron (UCN) target.



New field of physics unique to TRIUMF

Surface Physics

UCN

λ~100 Å

Ge(Hp)

detector

convertor

γ(477keV)

E. Korobkina, NCSU

thermal neutron

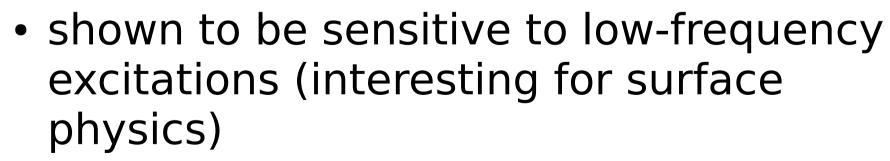
UCN

 Many ideas to use UCN to study 10 nm thin surface films

- (n,gamma)

- UCN loss measurements

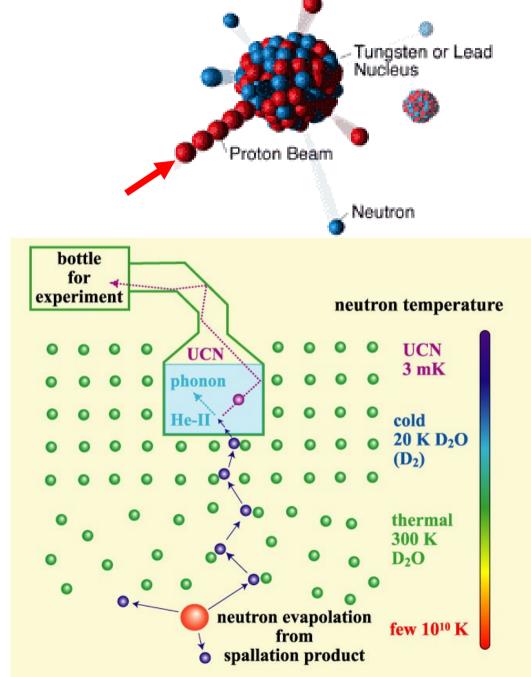
- n scattering
- reflectometry
 - polarization for magnetic films



In all cases, lack of UCN worldwide is the problem

How to make UCN

- Liberate neutrons by proton-induced spallation.
- Moderate (thermalize) in cold (20 K) D₂O.
- Cold neutrons then "downscatter" to near zero energy (4 mK) in superfluid helium through phonon production.



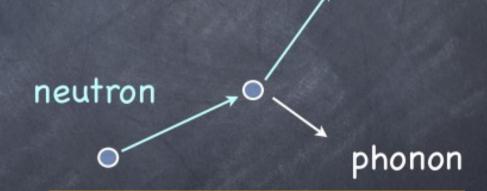
iah-eneray Proton

3.5 3.0 He-II 1.0 -0.5 Q (A')

M.R. Gibbs et al. (1999)

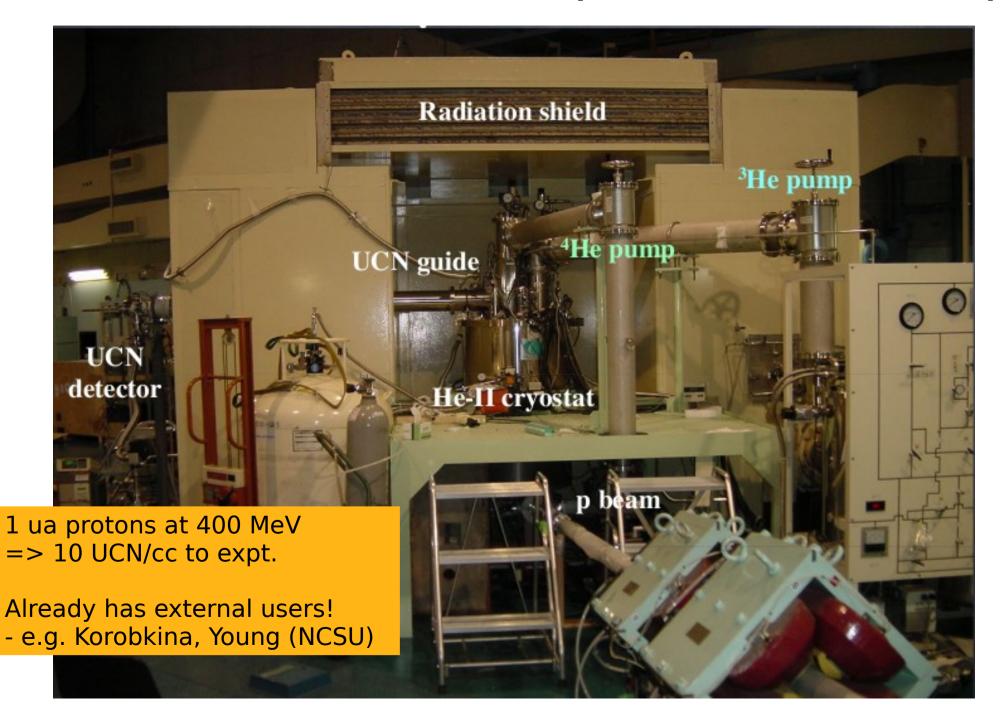
Superthermal UCN production in He-II

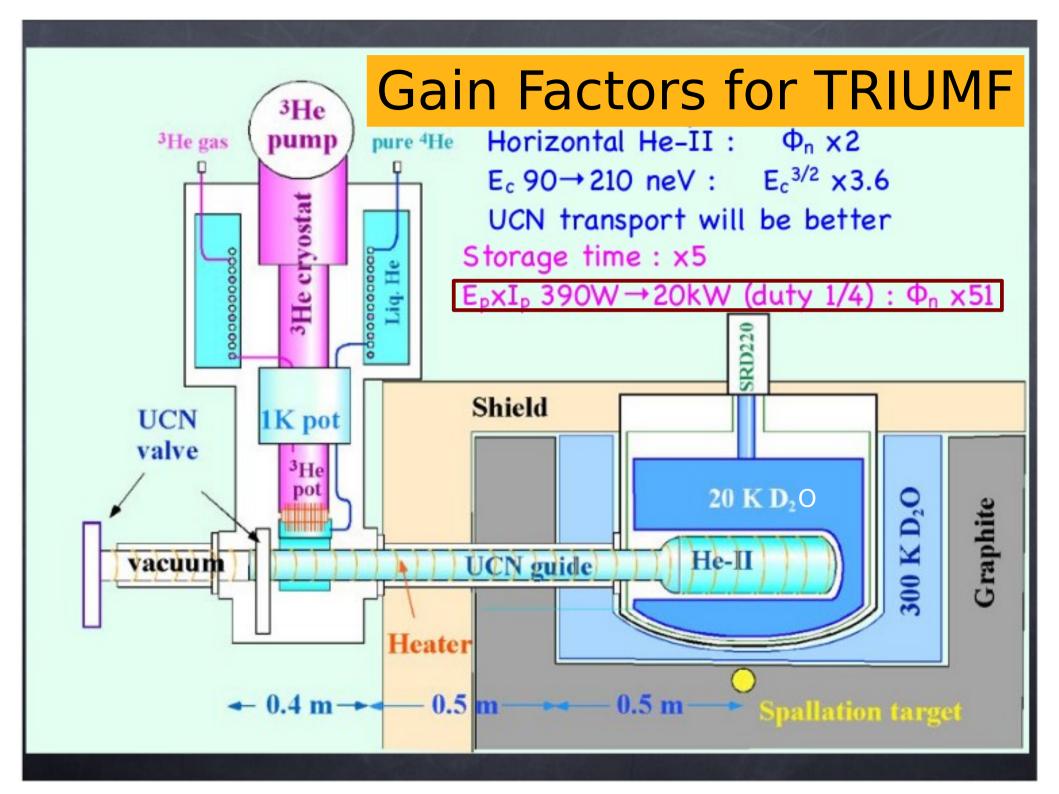
Coherent inelastic neutron scattering in He-II



Unique to TRIUMF!
Most other proposed sources
use solid deuterium instead
of superfluid helium

RCNP UCN Source (Masuda, et al)





Recent Progress on UCN at TRIUMF

- Aug. 2007 TUG working group (Canadians + Masuda), draft white paper.
- Sept. 2007 International UCN Workshop (+ world experts).
- Sept. 2007 Presentation to TRIUMF Board of Management
- Oct.-Nov. 2007 phone meetings relating to conceptual design
- Nov. 2007 ACOT

International Workshop: UCN Sources and Experiments

September 13-14, 2007 TRIUMF, Vancouver, Canada http://www.triumf.info/hosted/UCN

~25 speakers from all over the world ILL, FRM-II, NCSU, LANL, PSI, KEK, Mainz, ...

Supported by TRIUMF and TUNL

Speakers at TRIUMF UCN Workshop Sept. 13-14, 2007

- H. Abele, Heidelberg
- S. Baessler, Mainz/UVa
- L. Buchmann, TRIUMF
- M. Daum, PSI
- S. Gardner, U. Kentucky
- P. Geltenbort, ILL
- E. Gutsmiedl, Munich FRM-II
- R. Golub, NCSU
- B. Filippone, Caltech
- P. Huffman, NCSU
- T. Ito, LANL
- E. Korobkina, NCSU
- C.-Y. Liu, Indiana U.
- M. Makela, LANL

- J.W. Martin, U. Winnipeg
- Y. Masuda, KEK
- P. Mumm, NIST
- J. Nico, NIST
- J. Ng, TRIUMF
- M. Pospelov, U. Victoria / Perimeter Inst.
- J.-M. Poutissou, TRIUMF
- W.M. Snow, Indiana U.
- F. Wietfeldt, Tulane U.
- A. Young, NCSU
- G. Hampel, Mainz
- K. Hickerson, Caltech

talks available from http://www.triumf.info/hosted/UCN

Results of TRIUMF UCN Workshop

- Very strong statement from the international UCN community (particularly R. Golub) that a spallation He-II source should be pursued. Masuda clearly regarded as a leader in this field.
- TRIUMF would be an ideal venue for such a source.
- Many interesting physics experiments would be possible with the higher UCN densities achievable at this source.

TRIUMF UCN Source Schedule

- Prior to 2010, pursue development of new UCN cryostat for TRIUMF at RCNP, Osaka.
 - This would allow us to demonstrate all the gain factors from horizontal extraction, better UCN guides, etc. (aside from beam power)
- After 2010, begin construction of UCN source at TRIUMF (2010 = coincident with major reconstruction for ISAC+).

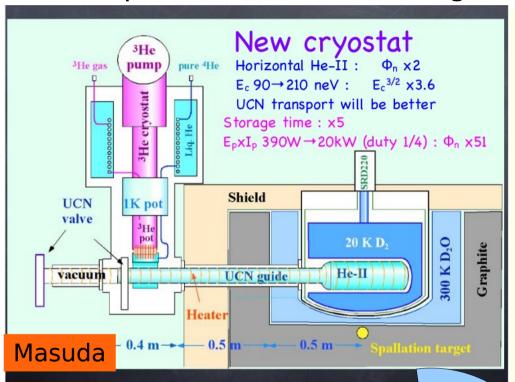
UCN Experimental Plan

- Conduct one initial high-profile, world's best, fundamental physics experiment (e.g. gravity, n-lifetime) (proposal to NSERC 2009)
- Simultaneous development of surface physics apparatus (2010)
- Developments towards future n-EDM experiment
- Other large-scale experiments (e.g. R-process, nnbar) (2015-)

Technical Issues for UCN

- Location.
- Beam sharing.
- Space (12m x 12m source + 12m x 6m expts)
- For precision physics experiments, it is highly advantageous to pulse the UCN source. E.g. RCNP uses 1 min beam on, 3 mins beam off. During beam off, UCN can be counted (or their decays, etc.)
 - pulsing at ion source incompatible with other users.
 - achieve pulsing by diverting beam to well-shielded dump using kicker.

Conceptual to technical design for the UCN source and experiment



Other Issues:

- radiation, remote handling
- cryogenics
- neutronics
- division of labor
 - Masuda, Golub very interested in cryogenics and neutronics
 - Acsion (private company in Pinawa, MB) interested in neutronics (MCNP)
 - need TRIUMF support to bring to fruition



Collaboration

- Masuda request to Japanese funding sources for 2.4 M\$ CAD over the next four years for UCN source plus EDM development (submitted Nov. 2007). The proposal explicitly outlines a plan to construct the source in Japan, and move it to TRIUMF.
- Intention to submit a CFI proposal from the Canadian groups in 2008 for the remaining infrastructure.
- Proposals to EEC and NSERC for the physics experiments.

Summary

- An opportunity exists to create the world's highest density UCN source at TRIUMF.
- A flagship physics experiment done this facility would be the world's best.
- UCN program is tied to the ISAC program in terms of both physics and facility.
- We are pursuing this unique and timely opportunity.

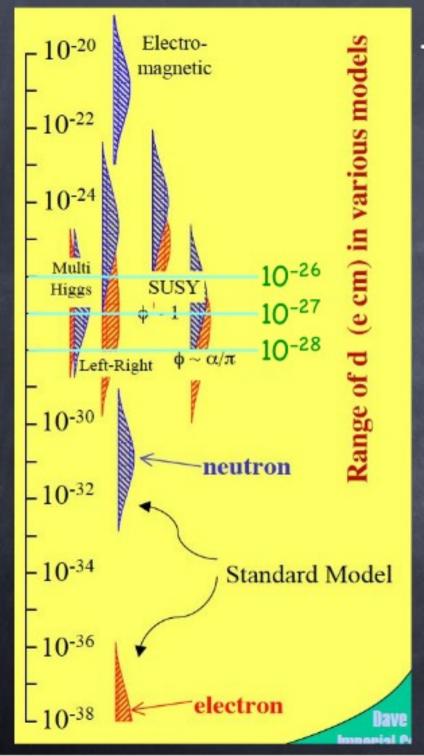
Back-ups

UCN Source Design and How We Would Achieve the World's Highest UCN Density

- Most other proposed sources use solid deuterium (ortho) at ~4K
 - ice quality
 - Fermi potential
 - para fraction, H-contamination
- We would use superfluid He-II (Masuda et al PRL)
 - lower specific production
 - NO loss mechanisms.

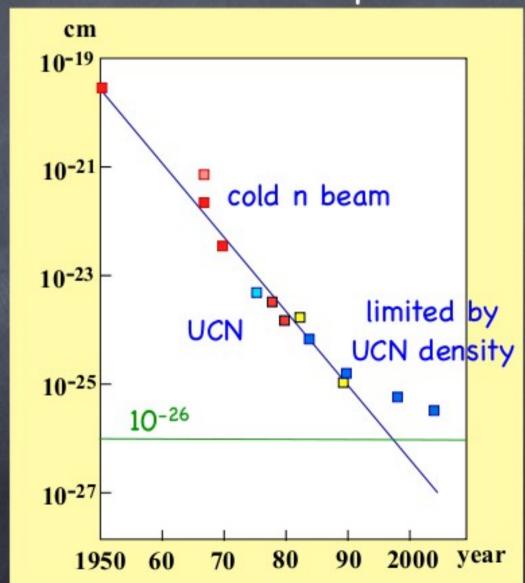
Cryogenics/Shielding Numbers from Masuda

- According to MC + estimates, for timeavg p-beam power of 12 kW in W target:
 - 0.45e12 n/cm2/s in He-II
 - 2.3 W in He-II
 - 30-60 W in 20 K D20
 - 1.7 kW in 300 K D20
- Masuda's current 3He pump can take 8 W.
- Clever arrangement of 208Pb can reduce gamma-heating of He-II even lower.
- Radiation #'s consistent/lower cf. LANL.



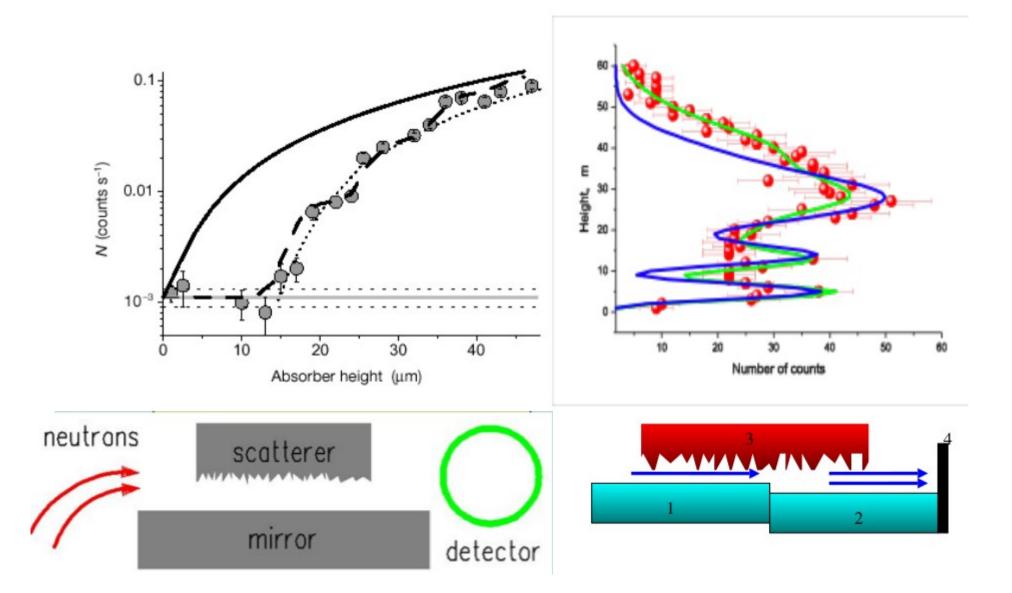
Theory EDM history

Experiment



UCN quantum states in gravity

• test of gravity at 10 um scale



Aug. 1 morning - plenary session, charge, and introduction of working groups 12:00 lunch 13:00 Welcome + Charge (Martin) (10+5) 13:15 UCN Sources Worldwide and for TRIUMF (Masuda) (45+10) 14:10 Photofission, (gamma,n) Sources and UCN (Behr) (10+10) 14:30 coffee (30) 15:00 SCRF joint session: Electron Linac Design (Koscielniak) (20+10) 15:30 UCN Infrastructure and Proton Hall Floorplan (Davis) (10+10) 15:50 Proton Hall Radiation Limits (Trudel) (10) 16:00 Discussion (Chair: Davis) (30) 16:30 tour of proton hall? (if desired) (Davis) (30) Aug. 2 09:00 Continued infrastructure Discussion (60) 10:00 UCN Physics Intro (Martin) (10+5) 10:15 UCN Beta Decay (Melconian) (20+10) 10:45 coffee (30) 11:15 n-EDM (Masuda and/or Hayden?) (20+10) 11:45 radioactive beams (Buchmann) (10+10) 12:05 lunch 13:15 gravity levels (Konaka) (5+10) 13:30 other physics (Martin) (15+10) 13:55 discussion of physics priorities and strategy (chair: Martin) (35) 14:30 begin writing more joint sessions draft Aug. 3 morning presentation Aug. 3

morning - presentations of results from the working groups.

Agenda

from townhall meeting (Aug. 07)