#### An Ultracold Neutron Source for TRIUMF

Jeff Martin (U. Winnipeg)

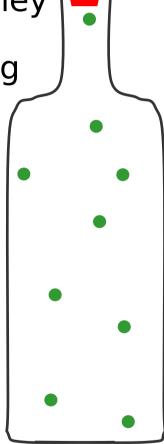
for Lothar Buchmann, Chuck Davis, Michael Gericke, Bob Golub, Akira Konaka, Larry Lee, Yasuhiro Masuda, Ania Micherdzinska, Shelley Page, Des Ramsay, Wim van Oers

(KEK, TRIUMF, NCSU, U. Manitoba, U. Winnipeg)

- 1. Introduction to Ultracold Neutrons (UCN)
- 2. UCN physics experiments
- 3. Source work at RCNP for TRIUMF
- 4. Plans for TRIUMF facility

## Ultracold Neutrons

- UCN are neutrons that are moving so slowly that they are totally reflected from surfaces of materials.
- So, they can be confined in material bottles for long periods of time.
- Typical parameters:
  - velocity < 8 m/s
  - temperature < 4 mK
  - kinetic energy < 300 neV</li>
- Interactions:
  - gravity: V = mgh (h < 3 m)
  - weak interaction (allows UCN to decay)
  - magnetic fields: V=- $\mu$ •B (100% polarization)
  - strong interaction
- Experiments at UCN sources are chronically limited by UCN density. TRIUMF has the potential to be a world leader in this regard.



#### UCN Source at TRIUMF would be a world-class facility

	Source type	$E_c$ and $\tau_s$	UCN density pucn(UCN/cm <sup>3</sup> )
TRIUMF 5 kW <sub>av</sub> proton	0.8K He-II	$\begin{array}{l} E_c = 210 \text{ neV} \\ \tau_s = 150 \text{ s} \end{array}$	1.8 x 10 <sup>4</sup> at experimental port
Grenoble 60MW reactor	0.5K He-II	$\begin{array}{l} E_c = 250 \text{ neV} \\ \tau_s = 150 \text{ s} \end{array}$	1000 in He-II
SNS cold neutron beam	0.3K He-II	$E_{c} = 134 \text{ neV} \ \tau_{s} = 500 \text{ s}$	430 in He-II
Munich 20MW reactor	SD <sub>2</sub>	$E_c = 250 \text{ neV}$	10 <sup>4</sup> in source
North Carolina 1 MW reactor	SD <sub>2</sub>	$E_c = 335 \text{ neV}$	1300 in source
PSI 12 kW <sub>av</sub> proton	SD <sub>2</sub>	$E_{c} = 250 \text{ neV} \ \tau_{s} = 888 \text{ s}$	2000 in source
Los Alamos 2.4 kW <sub>av</sub> proton	SD <sub>2</sub>	$E_{c} = 250 \text{ neV} \ \tau_{s} = 2.6 \text{ s}$	120 in source

#### The Most Successful UCN Experiments to Date

- n-EDM
- n-Lifetime
- Quantization of neutron energy levels in Earth's gravitation field
- Search for mirror neutrons

all of them were done at ILL, Grenoble

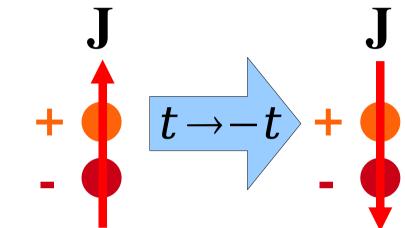
## Future UCN Experiments

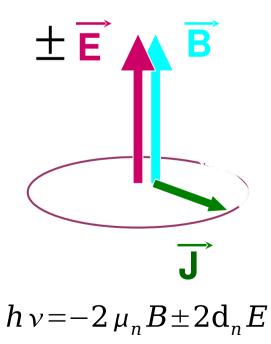
- EDM
  - in vacuo (PSI)
  - in 4He/3He (SNS)
  - ILL
  - KEK/JPARC
- Lifetime
  - gravito-magnetic bottle (LANL)
  - magnetic bottle in 4He (SNS)
- Gravity Levels
  - ILL with new 4He source

- β-Decay Correlations (A)
  - UCNA (LANL)
- r-process, free neutron target
- nnbar oscillations
- mirror neutrons
- surface physics
- physics of UCN production
  - e.g. O2

#### Neutron Electric Dipole Moment (n-EDM)

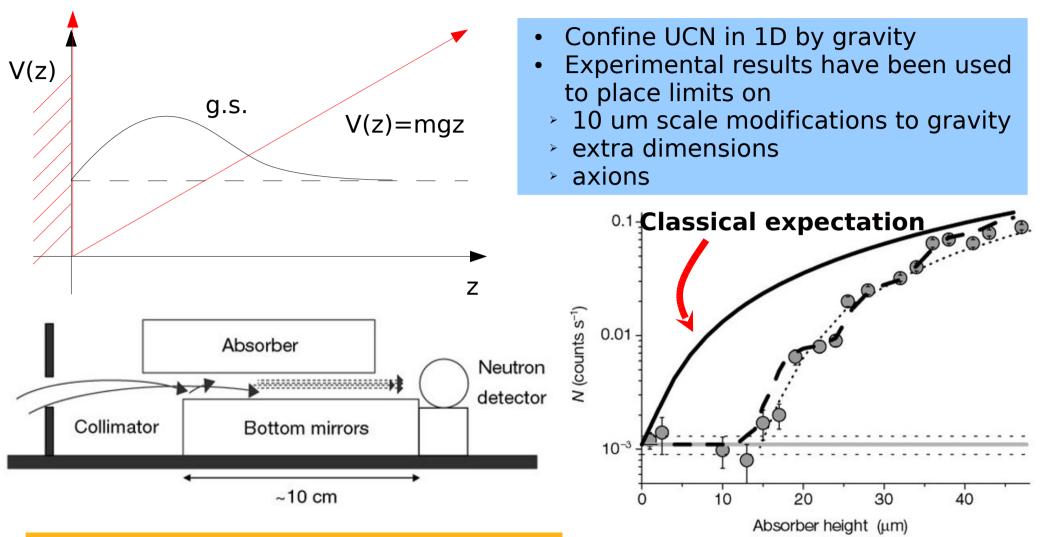
- Existence of EDM implies violation of Time Reversal Invariance
- CPT Theorem then implies violation of CP conservation





- Present Exp. Limit <  $0.63 \times 10^{-25}$  e-cm
- Standard Model value: 10-31 e-cm
- Supersymmetry or Multi-Higgs models can give 10<sup>5</sup>xSM
- Significant discovery potential with new high sensitivity *n*-EDM experiment

#### UCN Quantum States in Gravity

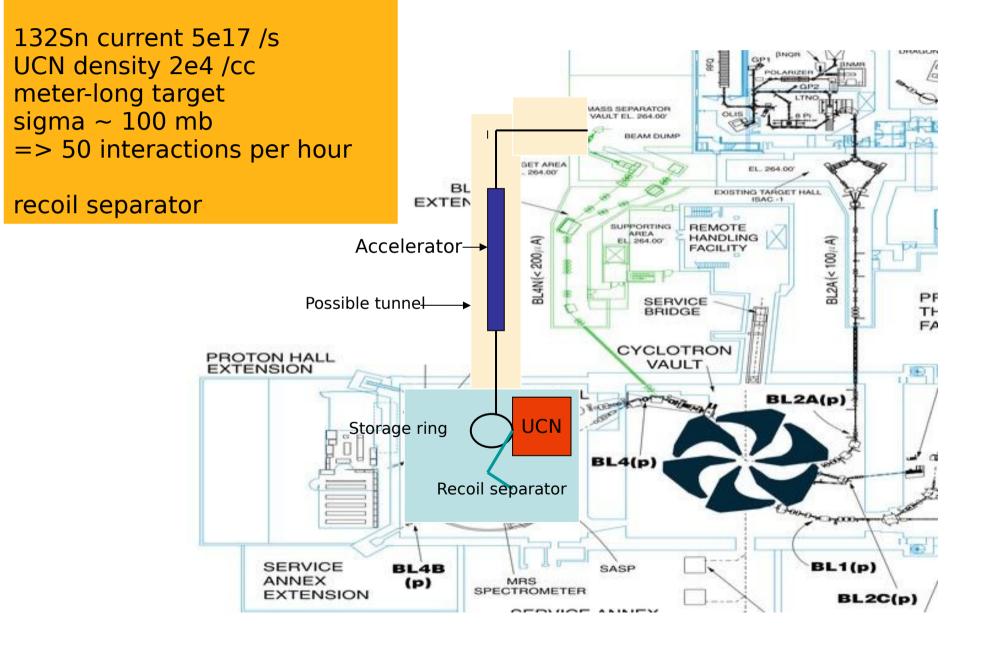


Further experiments:

- bottle the UCN to increase time the UCN is contact with the mirror.
- excite transitions between quantum states.
- increase purity of states by preselection.
- Goal: improve precision on energy of state and hence increase sensitivity to modifications to gravity.

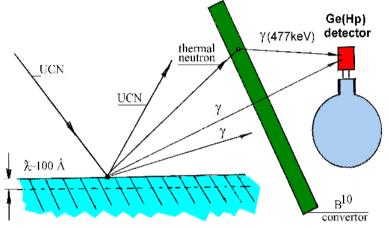
#### Measuring (n, y) cross sections of the r-process (Buchmann)

132Sn stored in ring interacts with free neutron (UCN) target.



#### Surface Physics

- Many ideas to use UCN to study 10 nm thin surface films
  - (n,gamma)
  - UCN loss measurements
  - n scattering
  - reflectometry



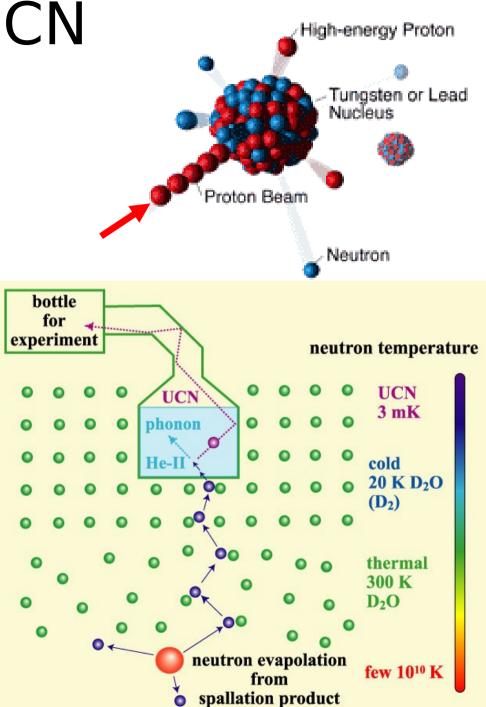
E. Korobkina, NCSU

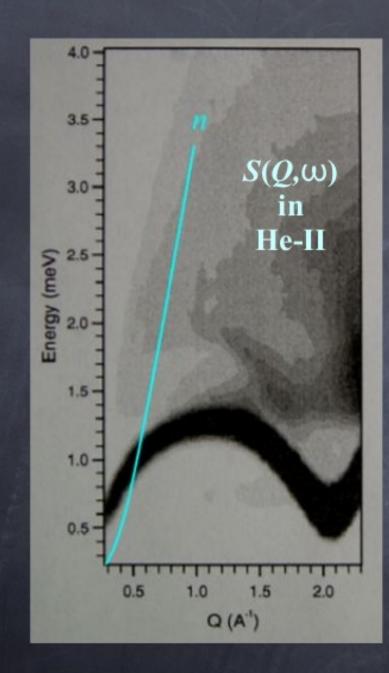
polarization for magnetic films

- shown to be sensitive to low-frequency excitations (interesting for surface physics)
- In all cases, lack of UCN worldwide is the problem

#### How to make UCN

- Liberate neutrons by proton-induced spallation.
- Moderate (thermalize) in cold (20 K) D<sub>2</sub>O.
- Cold neutrons then "downscatter" to near zero energy (4 mK) in superfluid helium through phonon production.





M.R. Gibbs et al. (1999)

## Superthermal UCN production in He-II

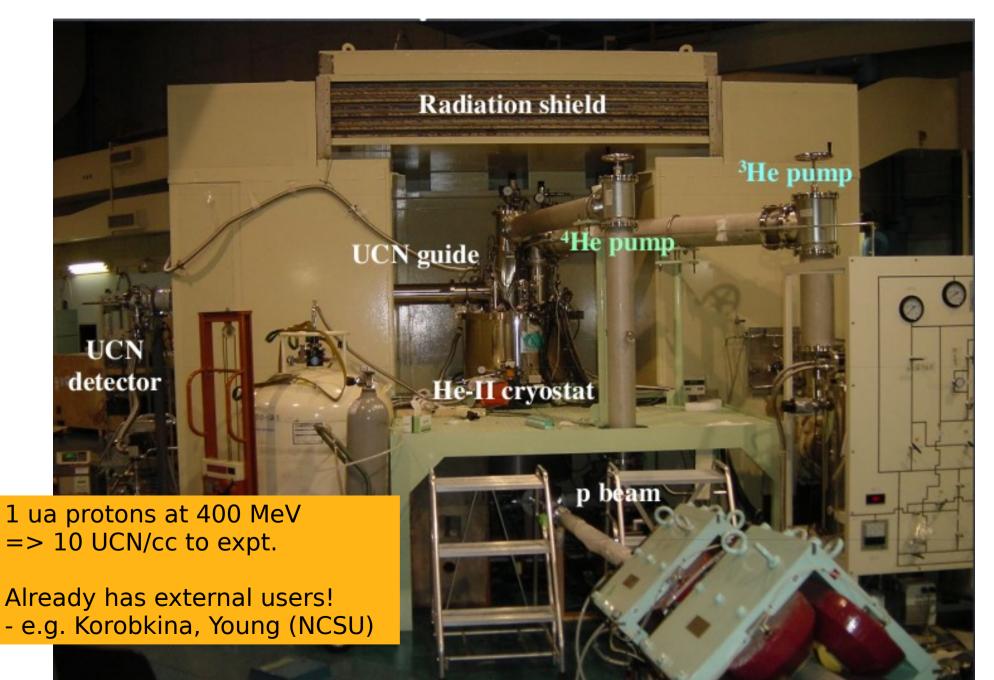
Coherent inelastic neutron scattering in He-II

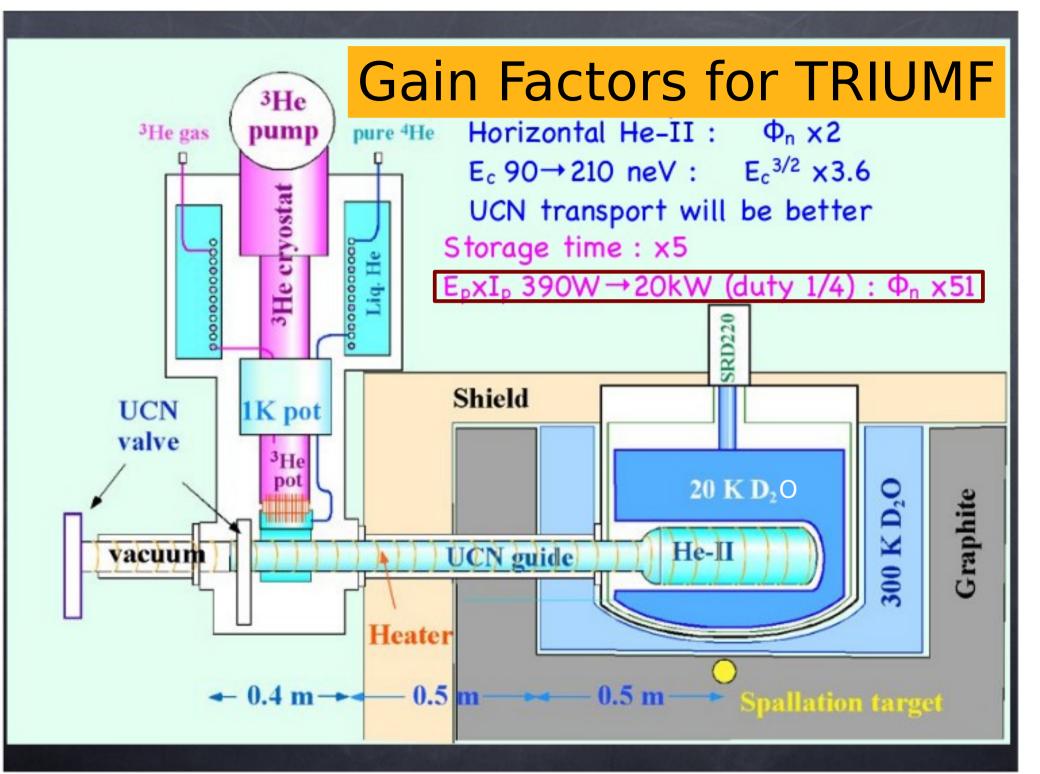
neutron

phonon

<u>Unique to TRIUMF!</u> Most other proposed sources use solid deuterium instead of superfluid helium

#### RCNP UCN Source (Masuda, et al)





#### Recent Progress on UCN at TRIUMF

- Aug. 2007 TUG working group (Canadians + Masuda), draft of white paper.
- Sept. 2007 International UCN Workshop (+ world experts).
- Sept. 2007 Presentation to Agency Committee on TRIUMF (ACT)
- Oct.-Nov. 2007 phone meetings relating to conceptual design
- Nov. 2007 ACOT

#### International Workshop: UCN Sources and Experiments

September 13-14, 2007 TRIUMF, Vancouver, Canada http://www.triumf.info/hosted/UCN

~25 speakers from all over the world ILL, FRM-II, NCSU, LANL, PSI, KEK, Mainz, ...

Supported by TRIUMF and TUNL

#### Speakers at TRIUMF UCN Workshop Sept. 13-14, 2007

- H. Abele, Heidelberg
- S. Baessler, Mainz/UVa
- L. Buchmann, TRIUMF
- M. Daum, PSI
- S. Gardner, U. Kentucky
- P. Geltenbort, ILL
- E. Gutsmiedl, Munich FRM-II
- R. Golub, NCSU
- B. Filippone, Caltech
- P. Huffman, NCSU
- T. Ito, LANL
- E. Korobkina, NCSU
- C.-Y. Liu, Indiana U.
- M. Makela, LANL

- J.W. Martin, U. Winnipeg
- Y. Masuda, KEK
- P. Mumm, NIST
- J. Nico, NIST
- J. Ng, TRIUMF
- M. Pospelov, U. Victoria / Perimeter Inst.
- J.-M. Poutissou, TRIUMF
- W.M. Snow, Indiana U.
- F. Wietfeldt, Tulane U.
- A. Young, NCSU
- G. Hampel, Mainz
- K. Hickerson, Caltech

talks available from http://www.triumf.info/hosted/UCN

#### Results of TRIUMF UCN Workshop

- Very strong statement from the international UCN community (particularly R. Golub) that a spallation He-II source should be pursued. Masuda clearly regarded as a leader in this field.
- TRIUMF would be an ideal venue for such a source.
- Many interesting physics experiments would be possible with the higher UCN densities achievable at this source.

#### TRIUMF UCN Source Schedule

- Prior to 2010, pursue development of new UCN cryostat for TRIUMF at RCNP, Osaka.
  - This would allow us to demonstrate all the gain factors from horizontal extraction, better UCN guides, etc. (aside from beam power)
- After 2010, begin construction of UCN source at TRIUMF (2010 = coincident with major reconstruction for ISAC 3).

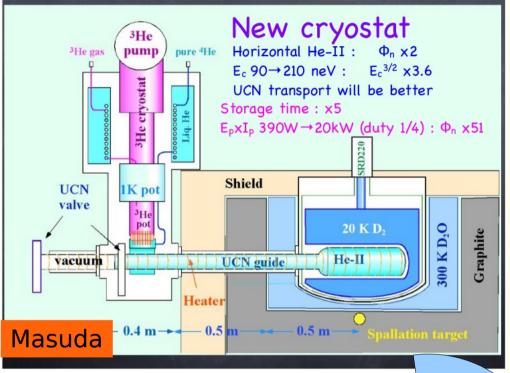
#### **UCN Experimental Plan**

- Conduct one initial high-profile, world's best, fundamental physics experiment (e.g. gravity, n-lifetime) (proposal to NSERC 2009)
- Simultaneous development of surface physics apparatus (2010)
- Developments towards future n-EDM experiments
- Other large-scale experiments (e.g. Rprocess, nnbar) (2015-)

#### **Technical Issues for UCN**

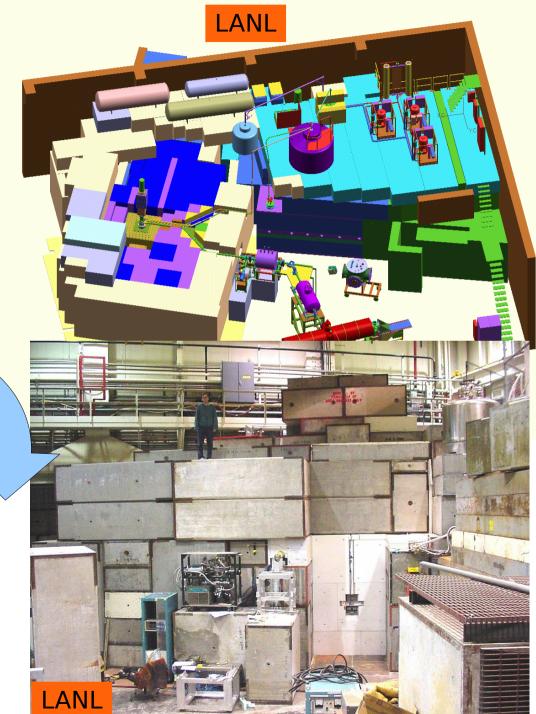
- Location.
- Beam sharing.
- Space (12m x 12m source + 12m x 6m expts)
- To carry out precision experiments, it is highly advantageous to pulse the UCN source. E.g. RCNP uses 1 min beam on, 3 mins beam off. During beam off, UCN can be counted (or their decays, etc.)
  - pulsing at ion source incompatible with other users.
  - achieve pulsing by diverting beam to well-shielded dump using kicker.

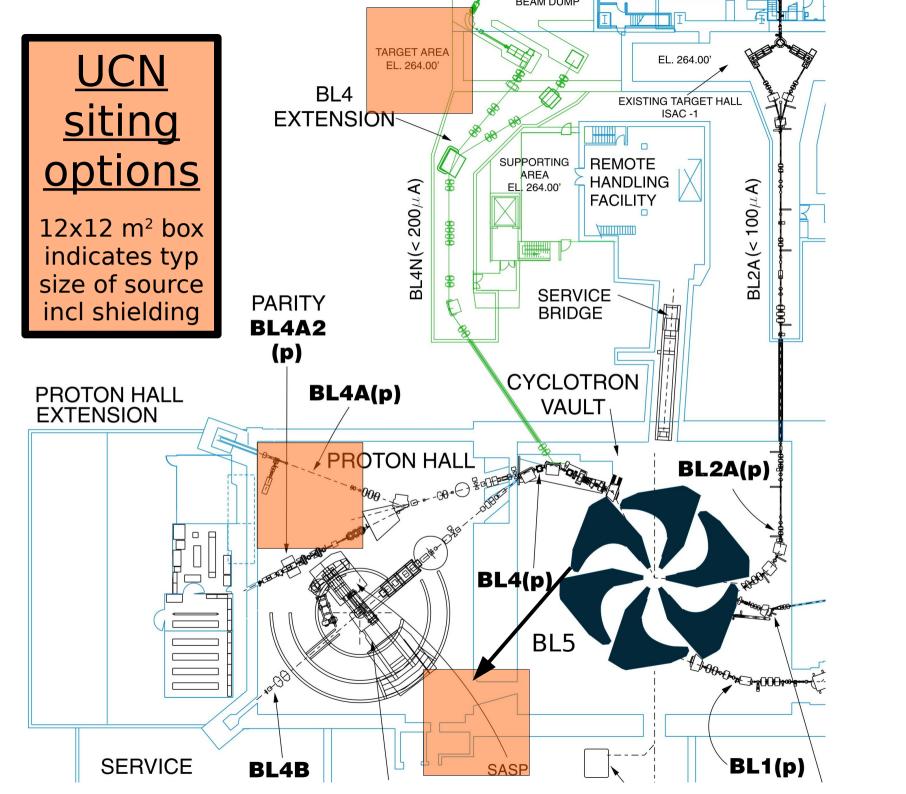
#### Conceptual to technical design for the UCN source and experiment

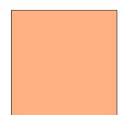


Other Issues:

- radiation, remote handling
- cryogenics
- neutronics
- division of tasks
  - Masuda, Golub very interested in cryogenics and neutronics
  - Acsion (private company in Pinawa, MB) interested in neutronics (MCNP)
  - need TRIUMF support to bring to fruition







#### Collaboration

- Masuda request to Japanese funding sources for 2.4 M\$ CAD over the next four years for UCN source plus EDM development (submitted Nov. 2007). The proposal explicitly outlines a plan to construct the source in Japan, and move it to TRIUMF.
- Intention to submit a CFI proposal from the Canadian groups in 2008 for the remaining infrastructure.
- Proposals to EEC and NSERC for the physics experiments.

#### Summary

- An opportunity exists to create the world's highest density UCN source at TRIUMF.
- A flagship physics experiment done this facility would be world's best.
- UCN program is tied to the ISAC program in terms of both physics and facility.
- We are pursuing this unique and timely opportunity.

#### Back-ups

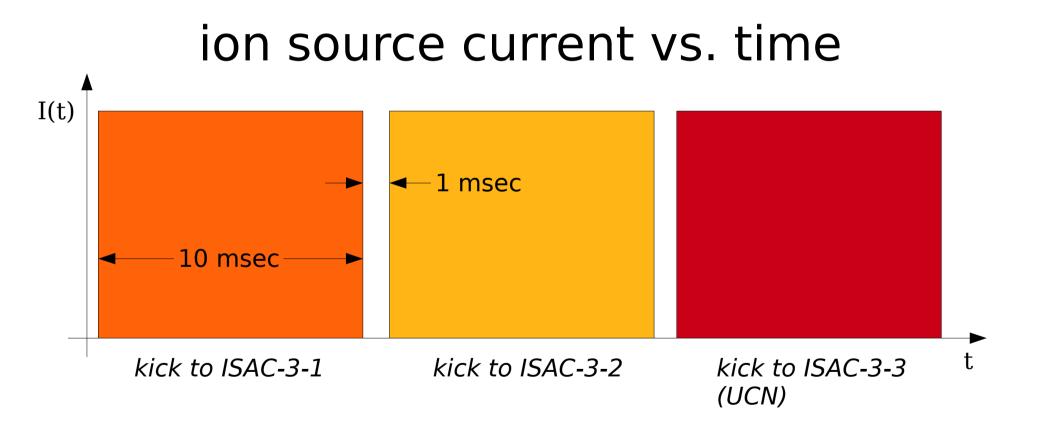
#### UCN Source Design and How We Would Achieve the World's Highest UCN Density

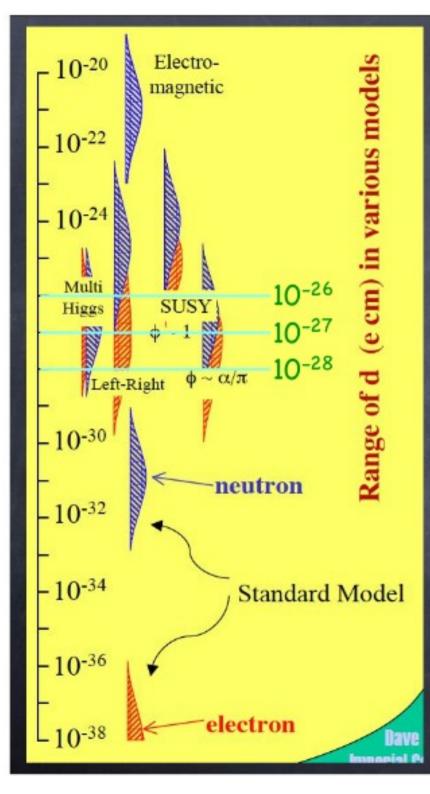
- Most other proposed sources use solid deuterium (ortho) at ~4K
  - ice quality
  - Fermi potential
  - para fraction, H-contamination
- We would use superfluid He-II (Masuda et al PRL)
  - lower specific production
  - NO loss mechanisms.

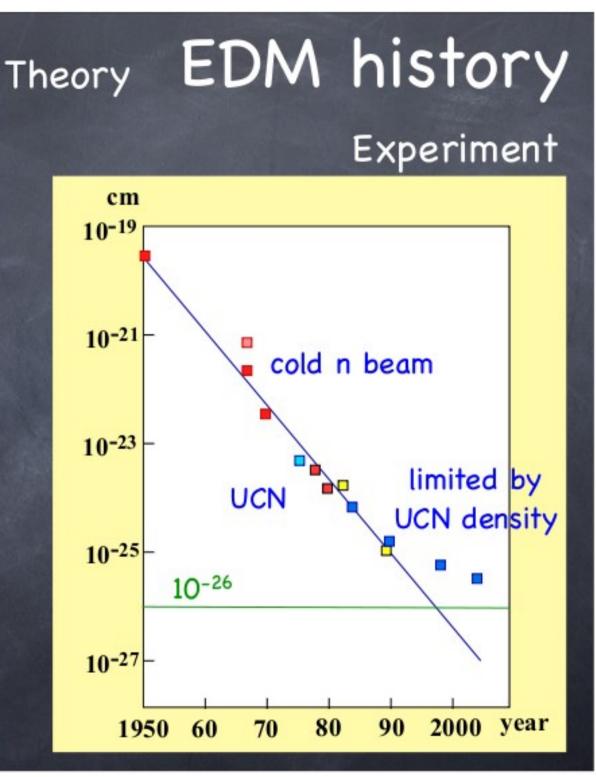
#### Cryogenics/Shielding Numbers from Masuda

- According to MC + estimates, for timeavg p-beam power of 12 kW in W target:
  - 0.45e12 n/cm2/s in He-II
  - 2.3 W in He-II
  - 30-60 W in 20 K D20
  - 1.7 kW in 300 K D20
- Masuda's current 3He pump can take 8 W.
- Clever arrangement of 208Pb can reduce gamma-heating of He-II even lower.
- Radiation #'s consistent/lower cf. LANL.

simultaneous operation with ISAC-3 by decoupling on kHz scale with kicker/ion source manipulation. Advantageous for ISAC-3: run all three targets simultaneously.

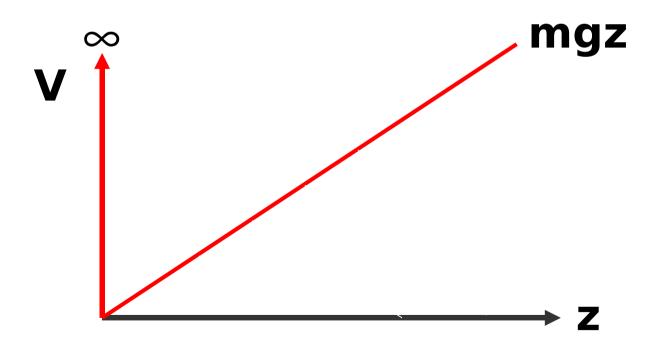




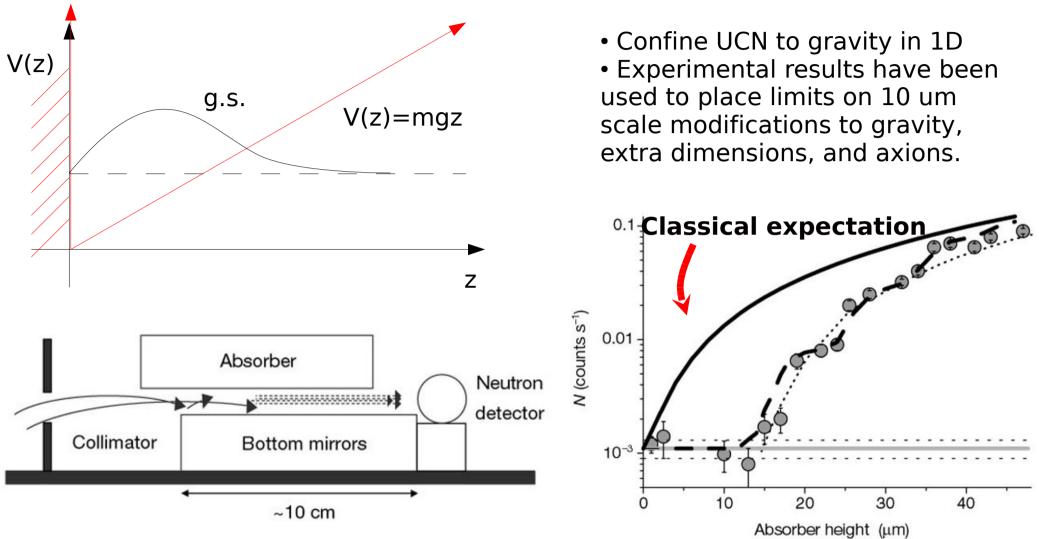


# Quantum States in Gravity Field

#### 1-d Schrodinger potential problem

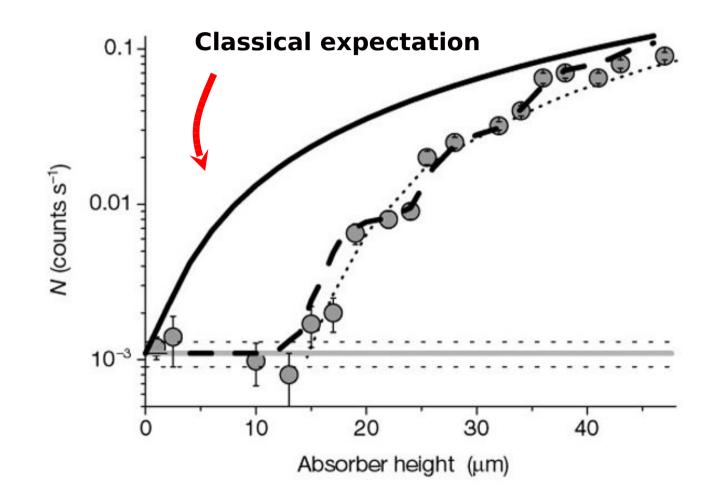


#### **UCN Quantum States in Gravity**



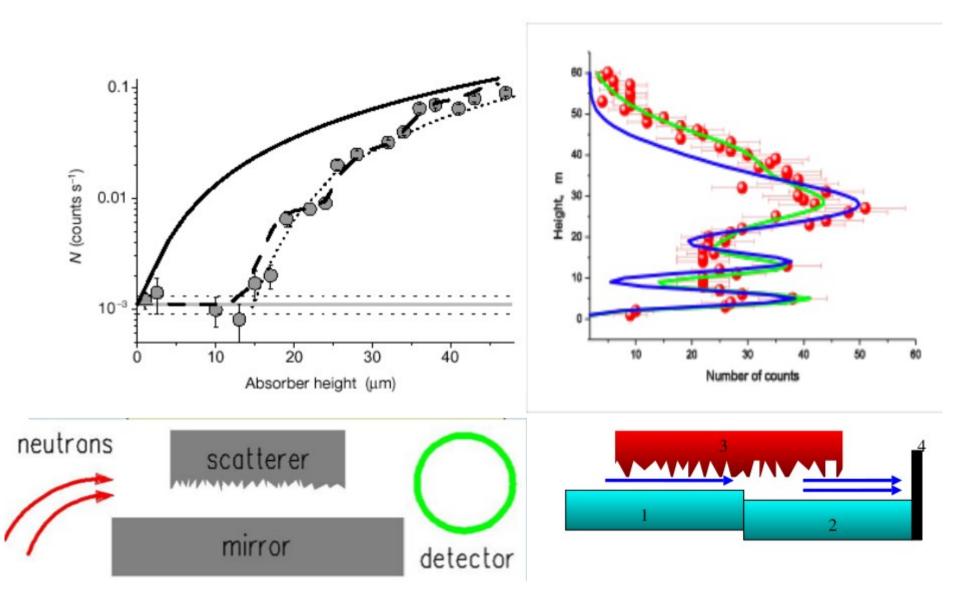
Suggestions for further experiments:

- bottle the UCN to increase time the UCN is contact with the mirror.
- excite transitions between quantum states.
- both these suggestions improve precision on energy of state and hence increase sensitivity to modifications to gravity.



# UCN quantum states in gravity

test of gravity at 10 um scale



Aug. 1

morning - plenary session, charge, and introduction of working groups

12:00 lunch

13:00 Welcome + Charge (Martin) (10+5)
13:15 UCN Sources Worldwide and for TRIUMF (Masuda) (45+10)
14:10 Photofission, (gamma,n) Sources and UCN (Behr) (10+10)
14:30 coffee (30)
15:00 SCRF joint session: Electron Linac Design (Koscielniak) (20+10)
15:30 UCN Infrastructure and Proton Hall Floorplan (Davis) (10+10)
15:50 Proton Hall Radiation Limits (Trudel) (10)
16:00 Discussion (Chair: Davis) (30)
16:30 tour of proton hall? (if desired) (Davis) (30)

Aug. 2

09:00 Continued infrastructure Discussion (60) 10:00 UCN Physics Intro (Martin) (10+5) 10:15 UCN Beta Decay (Melconian) (20+10) 10:45 coffee (30) 11:15 n-EDM (Masuda and/or Hayden?) (20+10) 11:45 radioactive beams (Buchmann) (10+10)

12:05 lunch

- 13:15 gravity levels (Konaka) (5+10)
- 13:30 other physics (Martin) (15+10)
- 13:55 discussion of physics priorities and strategy (chair: Martin) (35)
- 14:30 begin writing

more joint sessions

draft Aug. 3 morning presentation

Aug. 3

morning - presentations of results from the working groups.

# Agenda

from townhall meeting (Aug. 07)