An Ultracold Neutron Source for TRIUMF

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for Lothar Buchmann, Chuck Davis, Michael Gericke, Bob Golub, Akira Konaka, Larry Lee, Yasuhiro Masuda, Ania Micherdzinska, Shelley Page, Des Ramsay, Wim van Oers

(KEK, TRIUMF, NCSU, U. Manitoba, U. Winnipeg)

- 1. Introduction to Ultracold Neutrons (UCN)
- 2. UCN physics experiments
- 3. Source work at RCNP for TRIUMF
- 4. TRIUMF facility
- 5. CFI, relationship, collaboration, KEK

Ultracold Neutrons

- UCN are neutrons that are moving so slowly that they are totally reflected from surfaces of materials.
- So, they can be confined in material bottles for long periods of time.
- Typical parameters:
 - velocity < 8 m/s</p>
 - temperature < 4 mK
 - kinetic energy < 300 neV
- Interactions:
 - gravity: V = mgh (h < 3 m)
 - weak interaction (allows UCN to decay)
 - magnetic fields: V=- μ •B (100% polarization)
 - strong interaction
- Experiments at UCN sources are chronically limited by UCN density. TRIUMF has the potential to be a world leader in this regard.

Gravity

- V = mgh
- For a neutron on the planet Earth: $m = 1 \text{ GeV/c}^2$, $g = 10 \text{ m/s}^2$, h = 3 m $\Rightarrow V = 300 \text{ neV}$
- Recall, $T_{UCN} < 300 \text{ neV}$
- Uses:
 - UCN gravity spectrometer
 - gravitational levels expt

$\begin{array}{l} \text{Weak Interaction} \\ n \rightarrow p + e^{-} + \nu + 782 \text{ keV} \end{array}$

- $\tau = 885.7 + 0.8 s$ (needed for BBN)
- a, B, A (+ tau needed to extract Vud)

b, D, R (are zero in SM)

Magnetic Interaction

• The neutron has a magnetic moment:

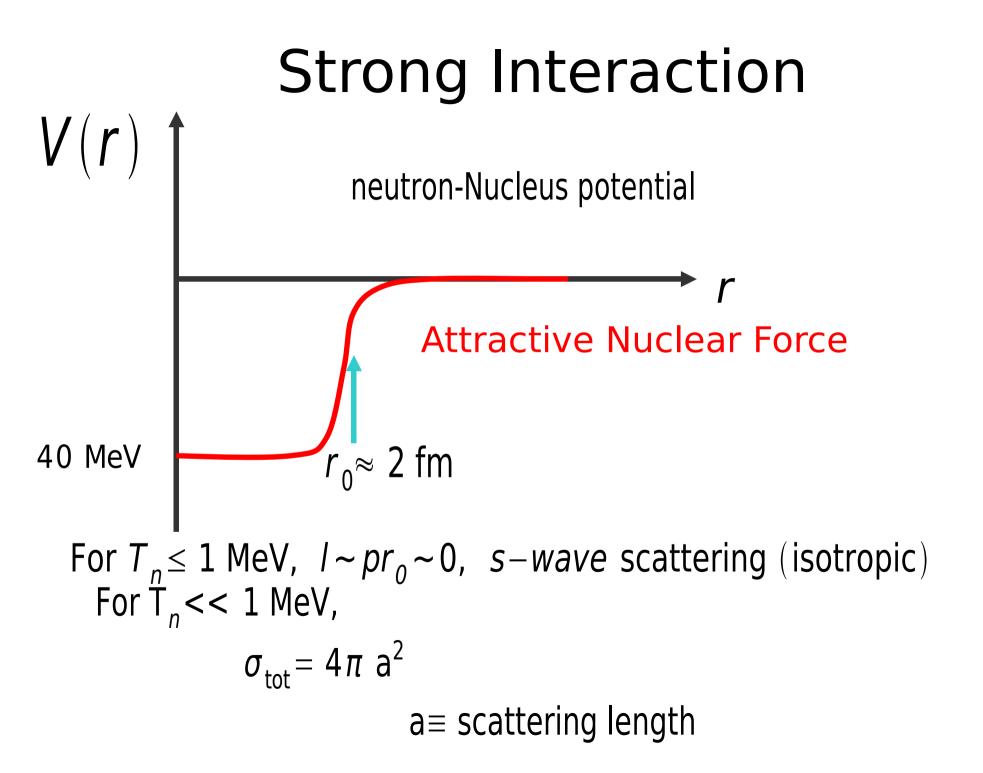
 μ = -1.9 μ_{N} = - 60 neV/T

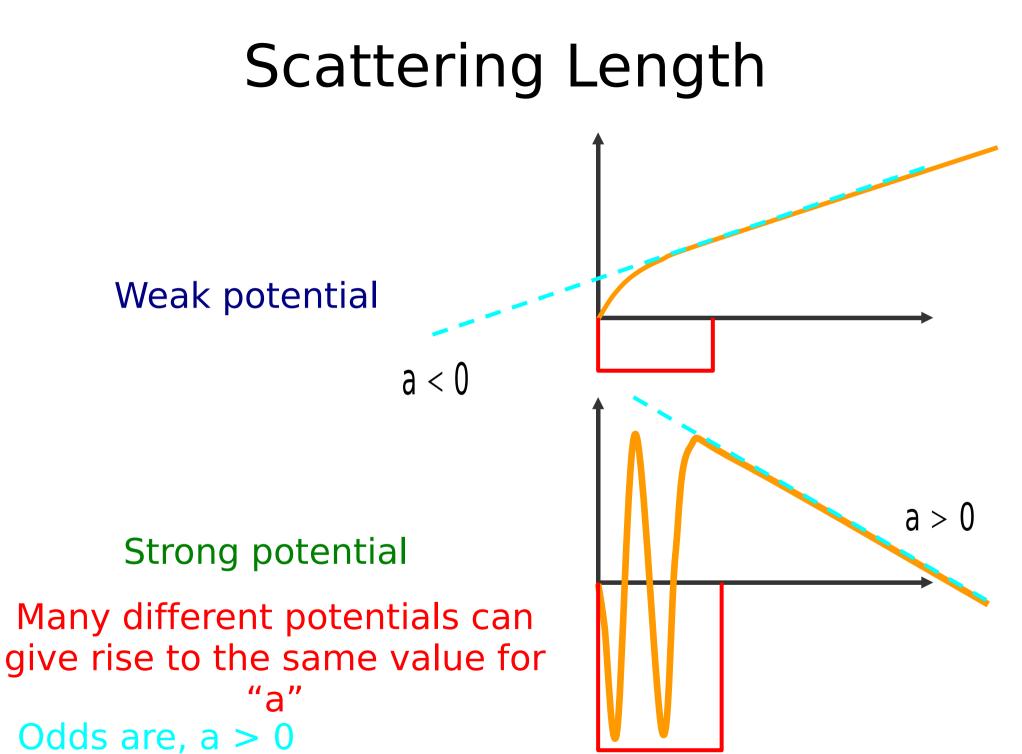
• The potential energy of a magnetic moment in a magnetic field is:

$$V = -\mu \bullet B$$

B = 7 T rejects one spin state of UCN (if adiabaticity obeyed)

=> UCN can be 100% polarized.

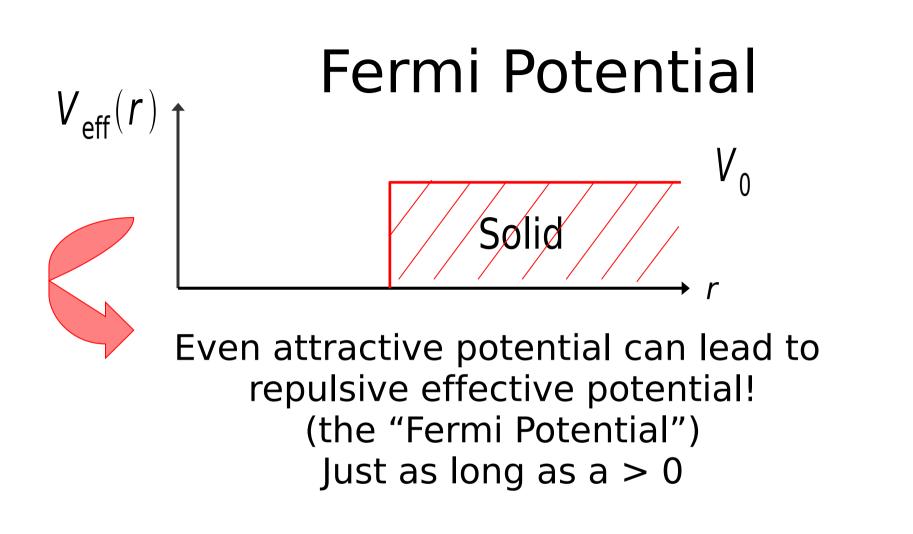




Fermi Potential

$$V_{eff}(r) = \frac{2\pi\hbar^2 a}{m} \delta(r)$$
$$V_{eff}(r) = \frac{2\pi\hbar^2}{m} \sum_{i} a_i \delta(r - r_i)$$

$$V_{eff}(r) = \frac{4\pi a \hbar^2}{2m} N_0 \int \frac{d^3 r'}{V} \delta(r - r') = \frac{2\pi a \hbar^2 n_0}{m} \theta(r \notin V) \equiv V_0 \theta(r \notin V)$$



Largest Fermi potential is for Nickel-58 (⁵⁸Ni) $V_0 = 335 \text{ neV}$

Physics Experiments Using UCN

• All previous experiments using UCN were chronically limited by UCN density.

UCN Source at TRIUMF would be a world-class facility

	Source type	E_c and τ_s	UCN density pucn(UCN/cm ³)
TRIUMF 5 kW _{av} proton	0.8K He-II	$\begin{array}{l} E_c = 210 \text{ neV} \\ \tau_s = 150 \text{ s} \end{array}$	1.8 x 10 ⁴ at experimental port
Grenoble 60MW reactor	0.5K He-II	$\begin{array}{l} E_c = 250 \text{ neV} \\ \tau_s = 150 \text{ s} \end{array}$	1000 in He-II
SNS cold neutron beam	0.3K He-II	$E_{c} = 134 \text{ neV} \ \tau_{s} = 500 \text{ s}$	430 in He-II
Munich 20MW reactor	SD ₂	$E_c = 250 \text{ neV}$	10 ⁴ in source
North Carolina 1 MW reactor	SD ₂	$E_c = 335 \text{ neV}$	1300 in source
PSI 12 kW _{av} proton	SD ₂	$E_{c} = 250 \text{ neV} \ \tau_{s} = 888 \text{ s}$	2000 in source
Los Alamos 2.4 kW _{av} proton	SD ₂	$E_{c} = 250 \text{ neV} \ \tau_{s} = 2.6 \text{ s}$	120 in source

The Most Successful UCN Experiments to Date

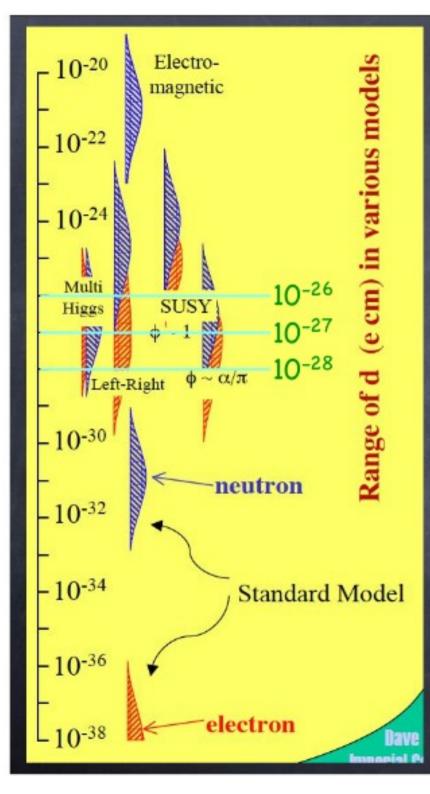
- EDM
- Lifetime
- Gravity Levels
- Mirror Neutrons

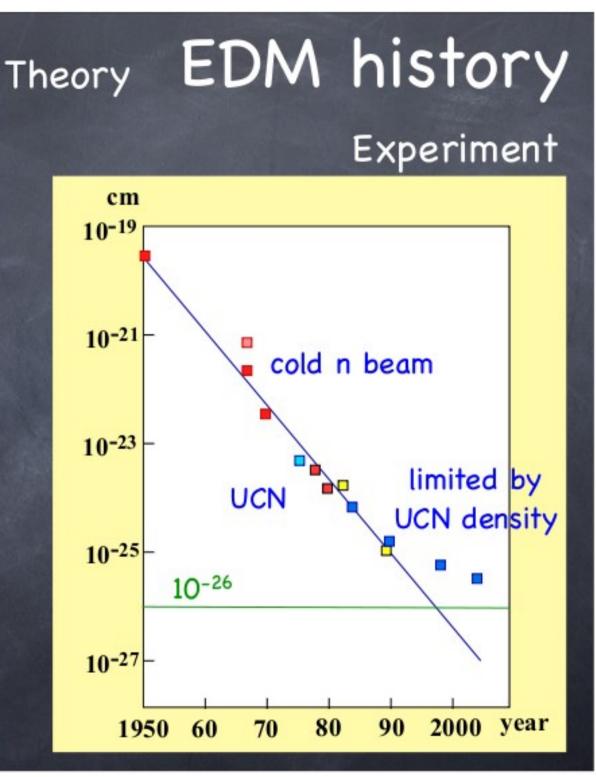
all of them were done at ILL, Grenoble

Future UCN Experiments

- EDM
 - in vacuo (PSI)
 - in 4He/3He (SNS)
 - ILL
 - KEK
- Lifetime
 - gravito-magnetic bottle (LANL)
 - magnetic bottle in 4He (SNS)
- Gravity Levels
 - ILL with new 4He source

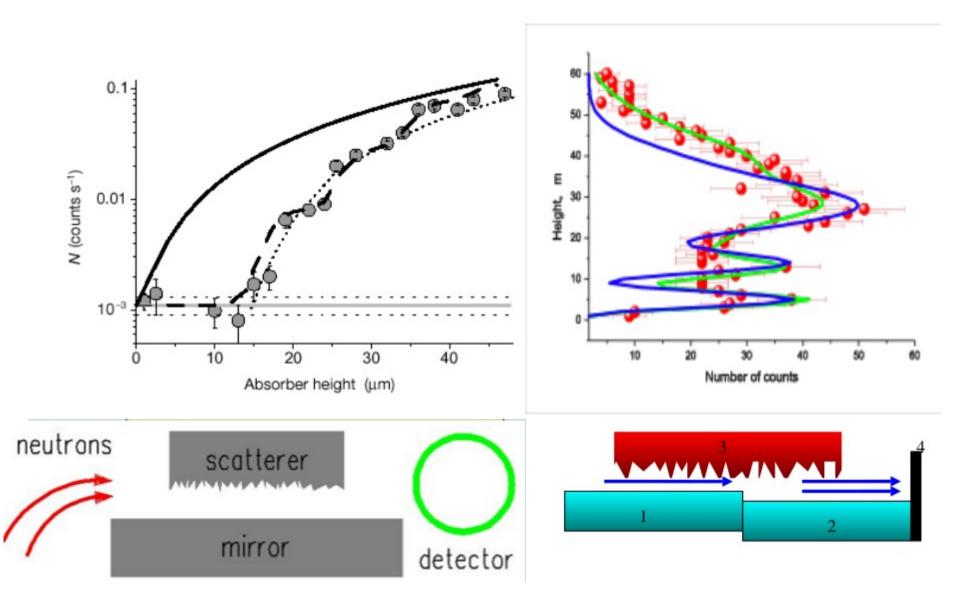
- β-Decay Correlations (A)
 - UCNA (LANL)
- r-process, free neutron target
- nnbar oscillations
- mirror neutrons
- surface physics
- physics of UCN production
 - e.g. 02





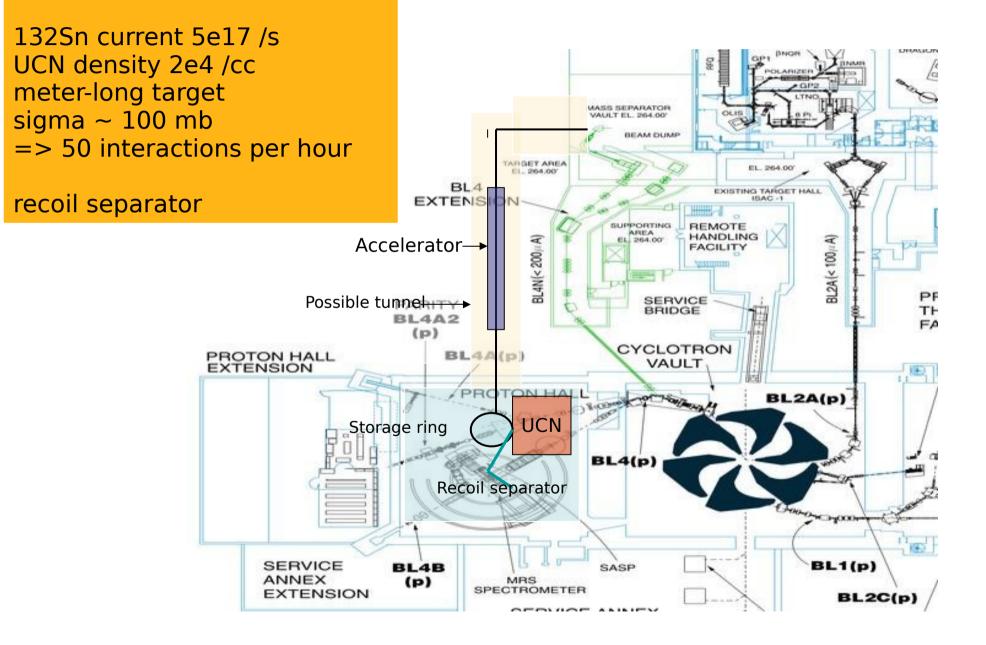
UCN quantum states in gravity

test of gravity at 10 um scale



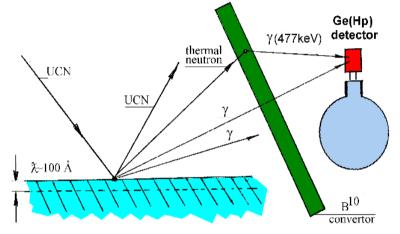
Measuring (n, y) cross sections of the r-process (Buchmann)

132Sn stored in ring interacts with free neutron (UCN) target.



Surface Physics

- Many ideas to use UCN to study 10 nm thin surface films
 - (n,gamma)
 - UCN loss measurements
 - n scattering
 - reflectometry



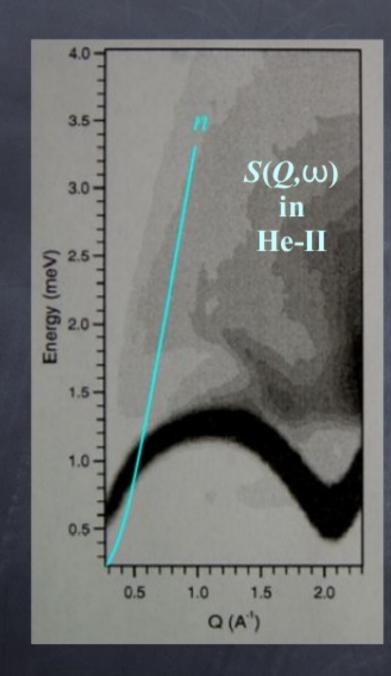
Korobkina, NCSU

polarization for magnetic films

- shown to be sensitive to low-frequency excitations (interesting for surface physics)
- In all cases, lack of UCN worldwide is the problem

UCN Source Design and How We Would Achieve the World's Highest UCN Density

- Most other proposed sources use solid deuterium (ortho) at ~4K
 - ice quality
 - material potential
 - para fraction, H-contamination
- We would use superfluid He-II (Masuda et al PRL)
 - lower specific production
 - NO loss mechanisms.



M.R. Gibbs et al. (1999)

Superthermal UCN production in He-II

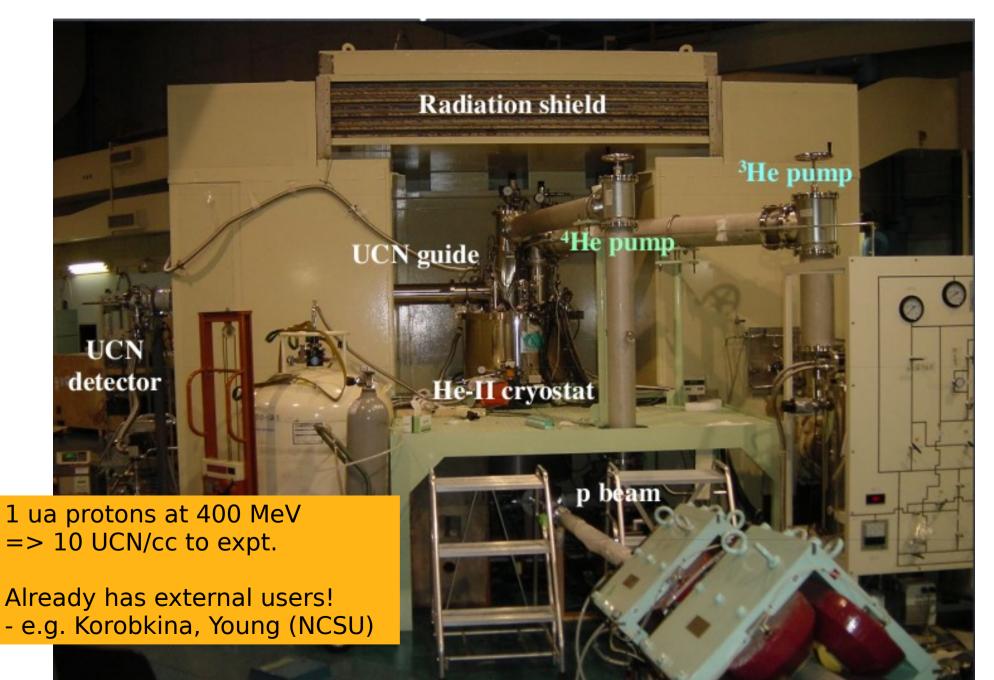
Coherent inelastic neutron scattering in He-II

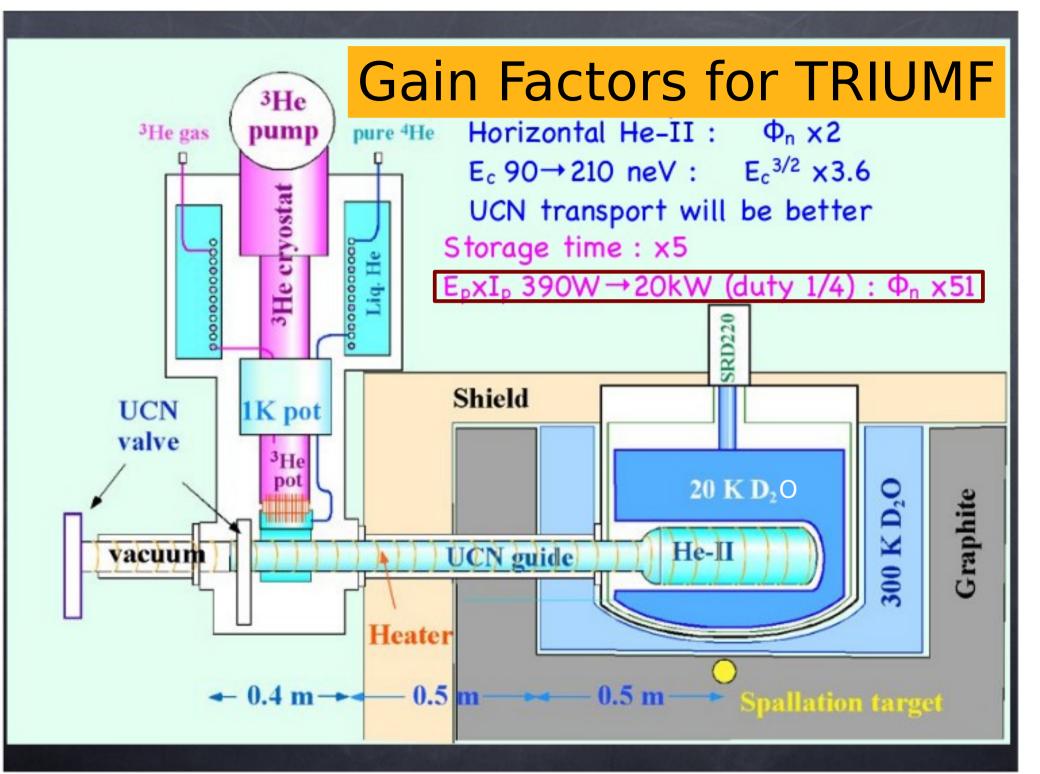
neutron

phonon

Born approximation $d^2\sigma/dQd\omega$ $= k_f/k_i a^2 S(Q,\omega)$ $= \sigma_{coh}/4\pi \cdot k_f/k_i \cdot S(Q,\omega)$

RCNP UCN Source (Masuda, et al)





Recent Progress on UCN at TRIUMF

- Aug. 2007 TUG working group (Canadians + Masuda), draft of white paper.
- Sept. 2007 International UCN Workshop (+ world experts).
- Sept. 2007 Presentation to Agency Committee on TRIUMF (ACT)
- Oct.-Nov. 2007 phone meetings relating to conceptual design
- Nov. 2007 ACOT

My old advertisement slide...

International Workshop: UCN Sources and Experiments

September 13-14, 2007 TRIUMF, Vancouver, Canada http://www.triumf.info/hosted/UCN

~25 speakers from all over the world ILL, FRM-II, NCSU, LANL, PSI, KEK, Mainz, ...

Supported by TRIUMF and TUNL

Speakers at TRIUMF UCN Workshop Sept. 13-14, 2007

H. Abele, Heidelberg

S. Baessler, Mainz/UVa

L. Buchmann, TRIUMF

M. Daum, PSI

S. Gardner, U. Kentucky

P. Geltenbort, ILL

E. Gutsmiedl, Munich FRM-II

R. Golub, NCSU

B. Filippone, Caltech

P. Huffman, NCSU

T. Ito, LANL

E. Korobkina, NCSU

C.-Y. Liu, Indiana U.

M. Makela, LANL

J.W. Martin, U. Winnipeg Y. Masuda, KEK P. Mumm, NIST J. Nico, NIST I. Ng, TRIUMF M. Pospelov, U. Victoria/Perimeter Inst. J.-M. Poutissou, TRIUMF W.M. Snow, Indiana U. F. Wietfeldt, Tulane U. A. Young, NCSU

G. Hampel, Mainz

K. Hickerson, Caltech

Results of Workshop

- Very strong statement from the international UCN community (particularly R. Golub) that a He-II source should be pursued. Masuda clearly regarded as a leader in this field.
- TRIUMF would be an ideal venue for such a source.
- Many interesting physics experiments would be possible with the higher UCN densities achievable at this source.

Schedule

- Prior to 2010, pursue development of new UCN cryostat for TRIUMF at RCNP, Osaka.
 - This would allow us to demonstrate all the gain factors from horizontal extraction, better UCN guides. (aside from beam power)
- After 2010, begin construction of UCN source at TRIUMF (2010 = coincident with major reconstruction for ISAC 3).

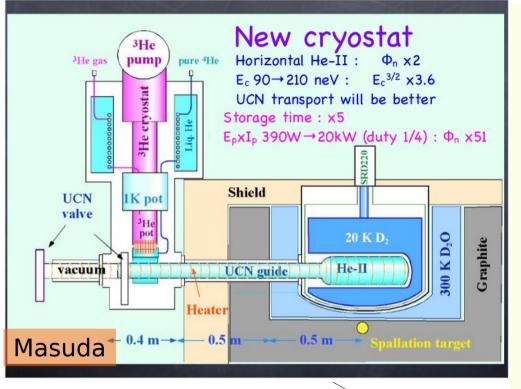
Experimental Plan

- Conduct one initial high-profile, world's best, fundamental physics experiment (likely either n-lifetime or gravity)
- Simultaneous development of surface physics apparatus
- Developments towards future n-EDM experiments
- Other large-scale experiments and transition to JPARC.

Technical Issues for UCN

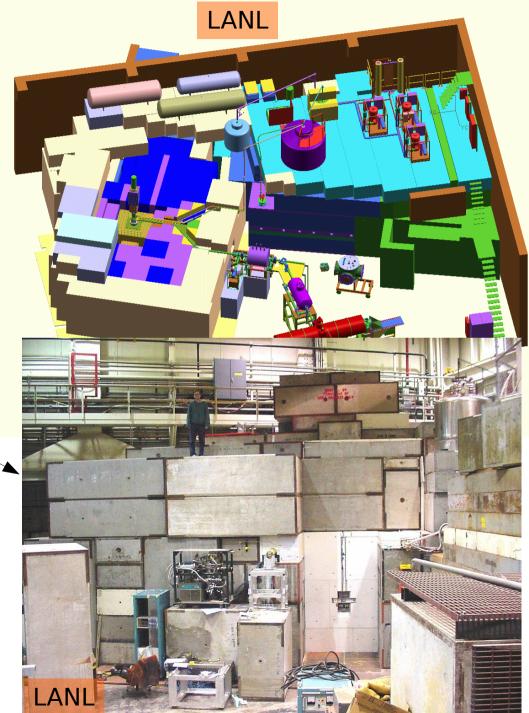
- Location.
- Beam sharing.
- Space (12m x 12m source + 12m x 6m expts)
- To carry out precision experiments, it is highly advantageous to pulse the UCN source. E.g. RCNP uses 1 min beam on, 3 mins beam off. During beam off, UCN can be counted (or their decays, etc.)
 - pulsing at ion source incompatible with other users.
 - achieve pulsing by diverting beam to well-shielded dump using kicker.

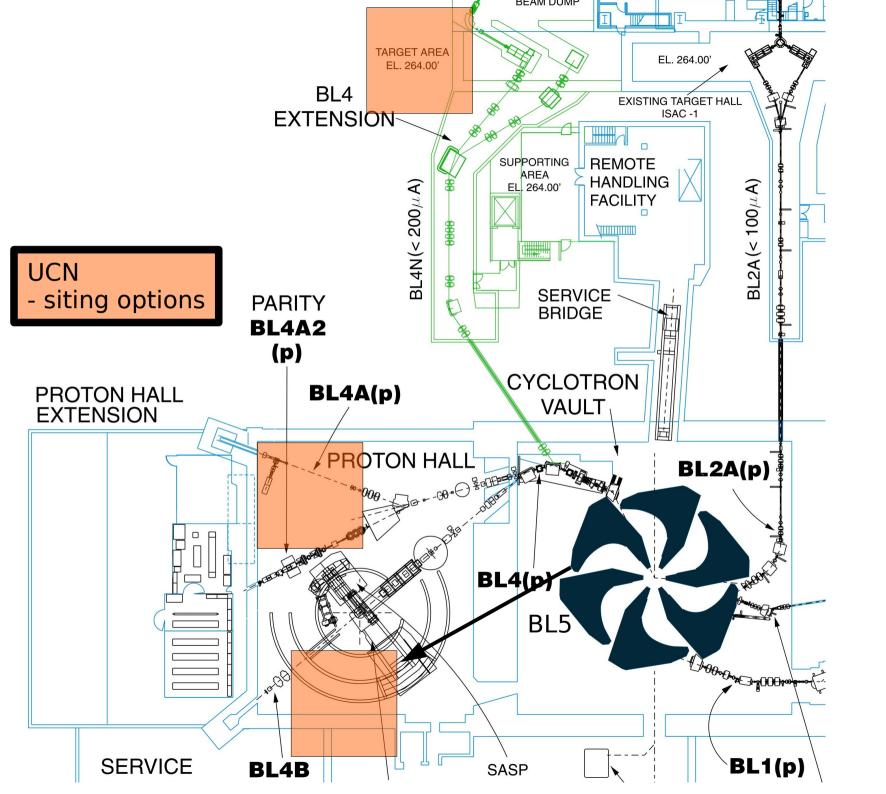
Conceptual to technical design for the UCN source and experiment



Other Issues:

- radiation, remote handling
- cryogenics
- neutronics
- division of tasks
 - Masuda, Golub, others? very interested in cryogenics and neutronics
 - need TRIUMF support to bring to fruition





Cost

- Cryostat, LHe cooling costs very wellunderstood (1.4 M\$ CAD)
- Masuda request to Japanese funding sources for 2.4 M\$ CAD over the next four years (the above plus EDM development) explicitly mentions TRIUMF. (submitted Nov. 2007)
- Interest in submitting a CFI proposal from the Canadian groups in 2008 for the remaining infrastructure.

Summary

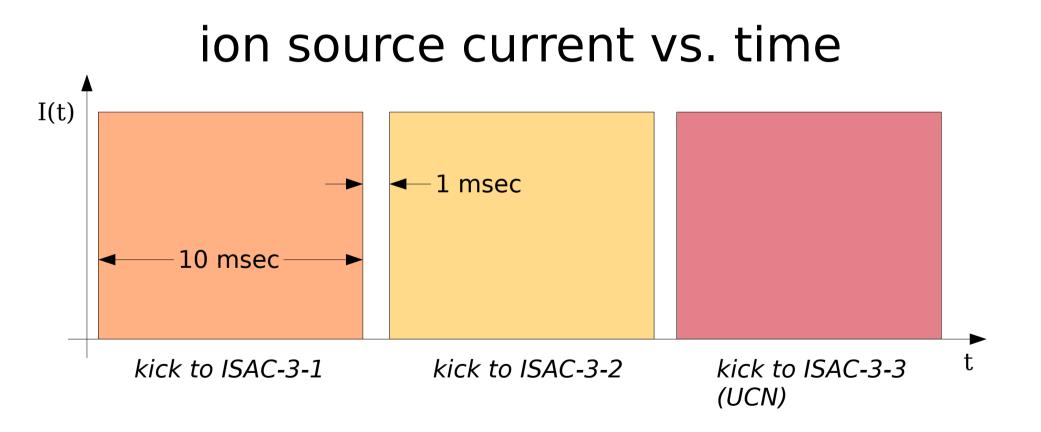
- An opportunity exists to create the world's highest density UCN source at TRIUMF
- A flagship physics experiment done this facility would be world's best
- The UCN facility would be tied to the ISAC-3 program in terms of both physics and facility
- We would like to pursue this unique and timely opportunity

Back-ups

Cryogenics/Shielding Numbers from Masuda

- According to MC + estimates, for timeavg p-beam power of 12 kW in W target:
 - 0.45e12 n/cm2/s in He-II
 - 2.3 W in He-II
 - 30-60 W in 20 K D20
 - 1.7 kW in 300 K D20
- Masuda's current 3He pump can take 8 W.
- Clever arrangement of 208Pb can reduce gamma-heating of He-II even lower.
- Radiation #'s consistent/lower cf. LANL.

simultaneous operation with ISAC-3 by decoupling on kHz scale with kicker/ion source manipulation. Advantageous for ISAC-3: run all three targets simultaneously.



Solutions Discussed for Technical Issues

- Location:
 - BL5 area
 - Advantages:
 - decoupled from ISAC 3.
 - Disadvantages:
 - new beam port must be constructed
 - beam line must not conflict with cyclotron probe extraction point
 - high-power dump required to achieve pulsing with kicker.
 - space in Proton Hall tighter likely requires excavation for dump.

Shameless advertising

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~25 speakers from all over the world ILL, FRM-II, NCSU, LANL, PSI, KEK, Mainz, ...

Supported by TRIUMF and TUNL

Aug. 1

morning - plenary session, charge, and introduction of working groups

12:00 lunch

13:00 Welcome + Charge (Martin) (10+5)
13:15 UCN Sources Worldwide and for TRIUMF (Masuda) (45+10)
14:10 Photofission, (gamma,n) Sources and UCN (Behr) (10+10)
14:30 coffee (30)
15:00 SCRF joint session: Electron Linac Design (Koscielniak) (20+10)
15:30 UCN Infrastructure and Proton Hall Floorplan (Davis) (10+10)
15:50 Proton Hall Radiation Limits (Trudel) (10)
16:00 Discussion (Chair: Davis) (30)
16:30 tour of proton hall? (if desired) (Davis) (30)

Aug. 2

09:00 Continued infrastructure Discussion (60) 10:00 UCN Physics Intro (Martin) (10+5) 10:15 UCN Beta Decay (Melconian) (20+10) 10:45 coffee (30) 11:15 n-EDM (Masuda and/or Hayden?) (20+10) 11:45 radioactive beams (Buchmann) (10+10)

12:05 lunch

- 13:15 gravity levels (Konaka) (5+10)
- 13:30 other physics (Martin) (15+10)
- 13:55 discussion of physics priorities and strategy (chair: Martin) (35)
- 14:30 begin writing

more joint sessions

draft Aug. 3 morning presentation

Aug. 3

morning - presentations of results from the working groups.

Agenda

from townhall meeting (Aug. 07)

Outline of White Paper

- Physics
 - prioritized and realistic, as much as possible
- UCN Source
 - proposed intensity at TRIUMF
 - world context and relevance
- Required Infrastructure
 - floorspace, shielding
 - duty cycle
 - He liquefier
- Required funding
 - CFI and collaboration, and international